

Higher spin gauge theory on fuzzy space(-times) from the IKKT matrix model

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Quantum Structure of Spacetime

Motivation

goal: quantum theory of spacetime & gravity

guidelines:

- simple
- gauge theory (Minkowski signature!)
- finite dof (per volume), pre-geometric (\rightarrow NCG ?)
- GR should arise as effective field theory (not starting point)

starting point:

Matrix Models as fundamental theories

- simple !
- describe dynamical NC / fuzzy spaces, **gauge theory**
- generic models of NCG: serious UV/IR mixing problem
preferred model: maximal SUSY: **IKKT model**
- preserves power of string theory, may avoid “landscape”

strategy: find good background space-time (solution)

→ typically issues with Lorentz invariance

way out: “covariant” quantum space → **higher spin** !

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outline:

- matrix models & matrix geometry
- *4D covariant quantum spaces*: fuzzy S_N^4, H_n^4
- cosmological space-times: $\mathcal{M}^{3,1}$ & BB!
- fluctuations \rightarrow *higher spin gauge theory*
- propagating modes, towards gravity on $\mathcal{M}^{3,1}$

HS, arXiv:1606.00769

M. Sperling, HS arXiv:1707.00885

HS, arXiv:1709.10480, arXiv:1710.11495

M. Sperling, HS arXiv:1806.05907

M. Sperling, HS arXiv:1901.xxxx

The IKKT model

IKKT or IIB model

Ishibashi, Kawai, Kitazawa, Tsuchiya 1996

$$S[X, \Psi] = -\text{Tr} \left([X^a, X^b][X^{a'}, X^{b'}] \eta_{aa'} \eta_{bb'} + m^2 X^a X_a + \bar{\Psi} \gamma_a [X^a, \Psi] \right)$$

$$X^a = X^{a\dagger} \in \text{Mat}(N, \mathbb{C}), \quad a = 0, \dots, 9, \quad N \text{ large}$$

gauge symmetry $X^a \rightarrow UX^aU^{-1}$, $SO(9, 1)$, SUSY

- quantized Schild action for IIB superstring
- reduction of $10D$ SYM to point, N large
- equations of motion:
 - $Y^a + m^2 Y^a = 0$, □ $\equiv \eta_{ab} [Y^a, [Y^b, \cdot]]$
- quantization: $Z = \int dX d\Psi e^{iS[X]}$

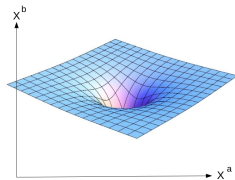
“matrix geometry” (\approx NC geometry):

- $S_E \sim \text{Tr}[X^a, X^b]^2 \Rightarrow$ config's with small $[X^a, X^b] \neq 0$ dominate

i.e. “almost-commutative” configurations

- \exists **quasi-coherent states** $|x\rangle$, minimize $\sum_a \langle x | \Delta X_a^2 | x \rangle$

$X^a \approx \text{diag.}$, spectrum $=: \mathcal{M} \subset \mathbb{R}^{10}$



NC branes embedded in target space \mathbb{R}^{10}

$$X^a \sim x^a : \quad \mathcal{M} \hookrightarrow \mathbb{R}^{10}$$

typical examples: **quantized Poisson manifolds**

- Moyal-Weyl quantum plane \mathbb{R}_θ^4 :

$$[X^a, X^b] = i\theta^{ab} \mathbf{1} \quad (\text{Heisenberg algebra})$$

quantized symplectic space (\mathbb{R}^4, ω)

admits translations $X^a \rightarrow X^a + c^a \mathbf{1}$, **no rotation invariance**

- fuzzy 2-sphere S_N^2

$$X_1^2 + X_2^2 + X_3^2 = R_N^2, \quad [X_i, X_j] = i\epsilon_{ijk} X_k$$

fully **covariant** under $SO(3)$

(Hoppe; Madore)

derivations:

$$[X^\mu, \phi] =: i\theta^{\mu\nu} \partial_\nu \phi$$

fluctuations in M.M. \rightarrow NC gauge theory

consider **fluctuations** around background X^a

$$Y^a = X^a + \mathcal{A}^a$$

$$[Y^\mu, \phi] = i\theta^{\mu\nu} D_\nu \phi, \quad D_\mu = \partial_\mu + i[A_\mu, \cdot]$$

$$[Y^\mu, Y^\nu] = i\theta^{\mu\nu} + i\theta^{\mu\mu'} \theta^{\nu\nu'} \underbrace{(\partial_{\mu'} A_{\nu'} - \partial_{\nu'} A_{\mu'} + [A_{\mu'}, A_{\nu'}])}_{F_{\mu'\nu'}}$$

$$S = \text{Tr}([Y^\mu, Y^\nu][Y_\mu, Y_\nu]) \sim \int F_{\mu\nu} F^{\mu\nu} + \text{c.}$$

\rightarrow YM gauge theory

\leftrightarrow dynamical geometry ("emergent gravity")

review: [H.S. arXiv:1003.4134](https://arxiv.org/abs/1003.4134)

covariant quantum spaces from equivariant bundles

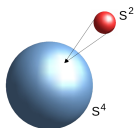
- issue: symplectic form ω breaks local (Lorentz/Euclid.) invar.
- way out: **quantized equivariant bundles** over \mathcal{M}^4
 - **higher spin** gauge theory
- example: **fuzzy four-sphere** S_N^4
 - Grosse-Klimcik-Presnajder; Castellino-Lee-Taylor; Ramgoolam; Kimura;
 - Karabail-Nair; Zhang-Hu 2001 (QHE!) ...
- noncompact H_n^4 → projection → cosmological space-time $\mathcal{M}_n^{3,1}$
- spin 2 modes on $\mathcal{M}_n^{3,1}$ $\xrightarrow{\text{conjecture}}$ gravity

(work in progress)

$\mathbb{C}P^3$ as S^2 bundle over S^4

Hopf map:

$$\begin{array}{ccc} \mathbb{C}P^3 & \ni & \psi \\ \downarrow & & \downarrow \\ S^4 & \ni & x^a = \psi^\dagger \gamma^a \psi \end{array}$$

... $SO(5)$ intertwiner, equivariant S^2 bundle $\mathbb{C}P^3$ is coadjoint orbit of $SU(4)$ or $SO(6) \supset SO(5)$

$$\mathbb{C}P^3 \cong \mathcal{O}[\Lambda] = \{g \cdot \Lambda \cdot g^{-1}; g \in SO(6)\} \subset \mathfrak{so}(6), \quad \Lambda = (0, 0, \lambda)$$

embedding functions

$$\begin{aligned} m^{ab} : \mathcal{O}[\Lambda] &\hookrightarrow \mathbb{R}^{15} \\ \xi &\mapsto m^{ab} = \text{tr}(\xi \Sigma^{ab}) \\ \xi &\mapsto x^a = \text{tr}(\xi \Sigma^{a6}) \end{aligned}$$

where $\xi = \psi\psi^\dagger$, $\Sigma^{ab} \in \mathfrak{so}(6)$, $\gamma^a = \Sigma^{a6}$

quantized (“fuzzy”) \mathcal{O}_Λ :

(Kirillov-Kostant-Souriau) symplectic form on $\mathcal{O}[\Lambda]$, Poisson structure

$$\{X^a, X^b\} = f^{abc} X^c$$

Λ ... dominant integral weight, \mathcal{H}_Λ ... h.w. irrep

$End(\mathcal{H}_\Lambda)$... quantized algebra of functions on \mathcal{O}_Λ

generators

$$[X^a, X^b] = if^{abc} X^c$$

covariant fuzzy four-sphere S_N^4

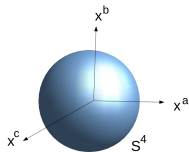
start with fuzzy $\mathbb{C}P^3$ given by $End(\mathcal{H}_\Lambda)$ for $\Lambda = (0, 0, N)$

quantized embedding function

$$X^a = \mathcal{M}^{a6} \sim x^a : \quad \mathbb{C}P^3 \rightarrow S^4 \subset \mathbb{R}^5$$

... quantized Hopf map !

(Grosse-Klimcik-Presnajder 1996)



relations:

$$\sum_a X_a^2 = R_N^2 \mathbf{1}, \quad R_N^2 \sim \frac{1}{4} N^2 \quad (\text{sphere})$$

$$\epsilon^{abcde} X_a X_b X_c X_d X_e = (N+2) R^2 \quad (\text{volume quantiz.})$$

$$[\mathcal{M}_{ab}, X_c] = i(\delta_{ac} X_b - \delta_{bc} X_a), \quad (\text{covariance})$$

\mathcal{M}_{ab} ... $so(5)$ generators acting on \mathcal{H}_N

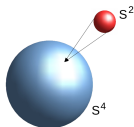
underlying fact: $(0, 0, N)$ of $so(6)$ is **irreducible** under $so(5)$

equivariant bundle structure

Poisson tensor $\theta^{\mu\nu}(x, \xi) \sim -i[X^\mu, X^\nu]$

local $SO(4)_x$ rotates fiber $\xi \in S^2$

averaging over fiber $\rightarrow [\theta^{\mu\nu}(x, \xi)]_0 = 0$, local $SO(4)$ preserved!



... 4D “covariant” quantum space

fields and harmonics on S_N^4

algebra of "functions":

$$\text{End}(\mathcal{H}_N) \cong \bigoplus_{s=0}^N C^s$$

$$C^s = \bigoplus_{n=0}^N (n, 2s) \ni \begin{array}{|c|c|c|c|} \hline & & & \\ \hline & & & \\ \hline & & & \\ \hline \end{array}$$

$(n, 0)$ modes = scalar functions on S^4 :

$$\phi(X) = \phi_{a_1 \dots a_n} X^{a_1} \dots X^{a_n} = \begin{array}{|c|c|c|c|} \hline & & & \\ \hline & & & \\ \hline & & & \\ \hline \end{array}$$

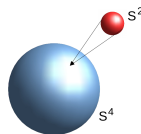
$(n, 2)$ modes = selfdual 2-forms on S^4

$$\phi_{bc}(X) \theta^{bc} = \phi_{a_1 \dots a_n b; c} X^{a_1} \dots X^{a_n} \theta^{bc} = \begin{array}{|c|c|c|c|} \hline & & & \\ \hline & & & \\ \hline & & & \\ \hline \end{array}$$

$\text{End}(\mathcal{H}) \cong$ fields on S^4 taking values in $\mathfrak{hs} = \oplus \begin{array}{|c|c|} \hline & \\ \hline & \\ \hline \end{array}$

higher spin modes = would-be KK modes on S^2

(local $SO(4)$ acts on S^2 fiber)



relations:

$$\begin{aligned}\mathcal{M}^{ab}\mathcal{M}^{ac} &= \frac{R^2}{\theta} P_T^{bc}, \\ \varepsilon_{abcde}\mathcal{M}^{ab}\mathcal{M}^{cd} &= 8\frac{R}{\theta} x^e,\end{aligned}$$

(\sim cf. bundle version of Joseph ideal ...)

relation with spin s fields: one-to-one map

$$\begin{aligned}
 \text{End}(\mathcal{H}_N) &\cong \bigoplus_{s=0}^N \mathcal{C}^s \cong \{\text{symmetric tensor field on } S^4\} \\
 \phi^{(s)} = \phi_{b_1 \dots b_s; c_1 \dots c_s}^{(s)}(x) \theta^{b_1 c_1} \dots \theta^{b_s c_s} &\mapsto \phi_{c_1 \dots c_s}^{(s)}(x) = \phi_{b_1 \dots b_s; c_1 \dots c_s}^{(s)} x^{b_1} \dots x^{b_s} \\
 \{x^{c_1}, \dots, \{x^{c_s}, \phi_{c_1 \dots c_s}^{(s)}(x)\} \dots\} &\leftarrow \phi_{c_1 \dots c_s}^{(s)}(x)
 \end{aligned}$$

... "symbol" of $\phi \in \mathcal{C}^s$

M. Sperling & HS, arXiv:1707.00885

$\mathcal{C}^s \cong$ symm., traceless, tang., div.-free rank s tensor field on S^4

$$\begin{aligned}
 \phi_{c_1 \dots c_s}(x) x^{c_i} &= 0, \\
 \phi_{c_1 \dots c_s}(x) g^{c_1 c_2} &= 0, \\
 \partial^{c_i} \phi_{c_1 \dots c_s}(x) &= 0.
 \end{aligned}$$

similarly:

cosmological quantum space-times $\mathcal{M}_n^{3,1}$:

- exactly homogeneous & isotropic, Big Bounce
- on-shell modes obtained
- starting point: fuzzy hyperboloid H_n^4

Euclidean fuzzy hyperboloid H_n^4 (=EAdS $_n^4$)

Hasebe arXiv:1207.1968

\mathcal{M}^{ab} ... hermitian generators of $\mathfrak{so}(4, 2)$,

$$[\mathcal{M}_{ab}, \mathcal{M}_{cd}] = i(\eta_{ac}\mathcal{M}_{bd} - \eta_{ad}\mathcal{M}_{bc} - \eta_{bc}\mathcal{M}_{ad} + \eta_{bd}\mathcal{M}_{ac}) .$$

choose “short” discrete unitary irreps \mathcal{H}_n $\eta^{ab} = \text{diag}(-1, 1, 1, 1, 1, -1)$ (“minireps”, doubletons)

special properties:

- irreps under $\mathfrak{so}(4, 1)$, multiplicities one, minimal oscillator rep.
- positive discrete spectrum

$$\text{spec}(\mathcal{M}^{05}) = \{E_0, E_0 + 1, \dots\}, \quad E_0 = 1 + \frac{n}{2}$$

lowest eigenspace is $n + 1$ -dim. irrep of $SU(2)_L$: fuzzy S_n^2

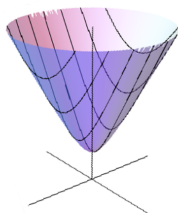
fuzzy hyperboloid H_n^4

def.

$$\begin{aligned} X^a &:= r\mathcal{M}^{a5}, & a = 0, \dots, 4 \\ [X^a, X^b] &= ir^2\mathcal{M}^{ab} =: i\Theta^{ab} \end{aligned}$$

5 hermitian generators $X^a = (X^a)^\dagger$ satisfy

$$\eta_{ab}X^aX^b = X^iX^i - X^0X^0 = -R^2\mathbf{1}, \quad R^2 = r^2(n^2 - 4)$$



one-sided hyperboloid in $\mathbb{R}^{1,4}$, covariant under $SO(4, 1)$

note: induced metric = Euclidean AdS^4

oscillator construction: 4 bosonic oscillators $[\psi_\alpha, \bar{\psi}^\beta] = \delta_\alpha^\beta$

$\mathcal{H}_n =$ suitable irrep in Fock space

Then

$$\mathcal{M}_{ab} = \bar{\psi} \Sigma_{ab} \psi, \quad \gamma_0 = \text{diag}(1, 1, -1, -1)$$

$$X^a = r \bar{\psi} \gamma^a \psi$$

$H_n^4 =$ quantized $\mathbb{C}P^{1,2} = S^2$ bundle over H^4 , selfdual $\theta^{\mu\nu}$

analogous to S_N^4

fuzzy "functions" on H_n^4 :

$$(End(\mathcal{H}_n) \rightarrow) \quad HS(\mathcal{H}_n) \cong \bigoplus_{s=0}^n \mathcal{C}^s = \int_{\mathbb{C}P^{1,2}} d\mu f(m) |m\rangle \langle m|$$

\approx fields on H^4 taking values in $\mathfrak{hs} = \bigoplus_s \begin{array}{|c|c|} \hline & \\ \hline & \\ \hline \end{array} \ni \mathcal{M}^{a_1 b_1} \dots \mathcal{M}^{a_s b_s}$

spin s sectors \mathcal{C}^s selected by **spin Casimir**

$$S^2 = \sum_{a < b \leq 4} [\mathcal{M}^{ab}, [\mathcal{M}_{ab}, \cdot]] + r^{-2} [X_a, [X^a, \cdot]],$$

can show:

$$S^2|_{\mathcal{C}^s} = 2s(s+1), \quad s = 0, 1, \dots, n$$

M. Sperling & H.S. 1806.05907

i.e. higher spin gauge theory, truncated at n

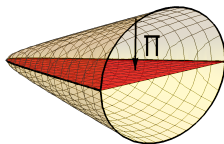
open FRW universe from H_n^4

HS arXiv:1710.11495

 $\mathcal{M}_n^{3,1} = H_n^4$ projected to $\mathbb{R}^{1,3}$ via

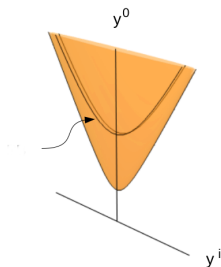
$$Y^\mu \sim y^\mu : \mathbb{C}P^{1,2} \rightarrow H^4 \subset \mathbb{R}^{1,4} \xrightarrow{\Pi} \mathbb{R}^{1,3} .$$

induced metric has Minkowski signature!

algebraically: $\mathcal{M}_n^{3,1}$ generated by

$$Y^\mu := X^\mu, \quad \text{for } \mu = 0, 1, 2, 3 \quad (\text{drop } X^4)$$

geometric properties:



- $SO(3, 1)$ manifest \Rightarrow foliation into $SO(3, 1)$ -invariant space-like 3-hyperboloids H_τ^3
- double-covered FRW space-time with hyperbolic ($k = -1$) spatial geometries

$$ds^2 = d\tau^2 - a(\tau)^2 d\Sigma^2,$$

$d\Sigma^2$... $SO(3, 1)$ -invariant metric on space-like H^3

"functions" on $\mathcal{M}_n^{3,1}$:

generated by $X^\mu = r\mathcal{M}^{\mu 5} \sim x^\mu$ and $T^\mu = \frac{1}{R}\mathcal{M}^{\mu 4} \sim t^\mu$

$$\begin{aligned}\{t^\mu, x^\nu\} &= \tilde{a}(\tau)\eta^{\mu\nu} \\ x_\mu t^\mu &= 0 \\ t_\mu t^\mu &= r^2(\tau)\end{aligned}$$

hence: t^μ ... space-like S^2 , derivations

self-duality on H^4 implies

$$\theta^{\mu\nu} = c(x^\mu t^\nu - x^\nu t^\mu) + b\epsilon^{\mu\nu\alpha\beta} x_\alpha t_\beta$$

→ expand modes as

$$\phi = \phi(x) + \phi_\mu(x)t^\mu + \phi_{\mu\nu}t^\mu t^\nu + \dots$$

"space-like gauge" (→ no ghosts!)

$\mathcal{M}^{3,1}$ realization in IKKT model:

background solution

$$T^\mu := \frac{1}{R} \mathcal{M}^{\mu 4}$$

satisfies

$$\square T^\mu = 3R^{-2} T^\mu, \quad \square = [T^\mu, [T_\mu, \cdot]]$$

- $[\square, \mathcal{S}^2] = 0$

⇒ higher-spin expansion $\phi = \phi(X) + \phi_\mu(X) T^\mu + \dots \in HS(\mathcal{H})$
on $\mathcal{M}^{3,1}$

- \square encodes effective FRW metric, asymptotically coasting

Big Bounce, initial $a(\tau) \sim \tau^{1/5}$ singularity

... work in progress

M. Sperling & HS

fluctuations & higher spin gauge theory

$$S[Y] = \text{Tr}(-[Y^a, Y^b][Y_a, Y_b] + m^2 Y^a Y_a) = S[U^{-1} Y U]$$

background solution: $S_N^4, H_n^4, \mathcal{M}_n^{3,1}$

add **fluctuations** $Y^a = \bar{Y}^a + \mathcal{A}^a,$

gauge trafos $\mathcal{A}^a \rightarrow [\Lambda, \mathcal{A}^a] + [\Lambda, \bar{Y}^a], \quad \Lambda \in \text{End}(\mathcal{H})$

expand action to second order in \mathcal{A}^a

$$S[Y] = S[\bar{Y}] + \frac{2}{g^2} \text{Tr} \mathcal{A}_a \left(\underbrace{\left(\square + \frac{1}{2} m^2 \right) \delta_b^a + 2[[\bar{Y}^a, \bar{Y}^b], \cdot]}_{\mathcal{D}^2} - \underbrace{[\bar{Y}^a, [\bar{Y}^b, \cdot]]}_{g.f.} \right) \mathcal{A}_b$$

- **fluctuations** \mathcal{A}_a describe gauge theory (NCFT) on \mathcal{M}
("open strings" ending on \mathcal{M})
- for S_N^4, H_n^4 : \mathcal{A}_a ... \mathfrak{hs} -valued gauge field, incl. spin 2

on S_N^4 and H_n^4 : background $\bar{Y}^a = X^a$

- $\mathcal{A}_a \in \text{End}(\mathcal{H}) \otimes (5)$

→ 4 indep. tangential fluctuation modes for each spin $s \leq n$

$$\begin{aligned} \mathcal{A}_a^{(1)} &= \bar{\partial}_a \phi^{(s)} \in \mathcal{C}^s, & \bar{\partial}^a &= x_b \{ \theta^{ba}, . \} \\ \mathcal{A}_a^{(2)} &= \theta^{ab} \bar{\partial}_b \phi^{(s)} = \{ x^a, \phi^{(s)} \} \\ \mathcal{A}_a^{(3)} &= \phi_a^{(s)}, & \phi^{(s)} &= \{ x^a, \phi_a^{(s)} \} \\ \mathcal{A}_a^{(4)} &= \theta^{ab} \phi_b^{(s)} \end{aligned}$$

+ radial mode $\mathcal{A}_a^{(r)} = x_a \phi^{(s)}$

- diagonalize → **eigenmodes** of \mathcal{D}^2

(details: M. Sperling & H.S. 1707.00885, 1806.05907)

all tangential modes are stable !

- radial modes are unstable on H_n^4

→ project to cosmological space-time $\mathcal{M}^{3,1}$

on $\mathcal{M}_n^{3,1}$: background $\bar{Y}^\mu = T^\mu$ M. Sperling and HS, work in progress

symmetry reduced to space-like $SO(3, 1)$

underlying $SO(4, 1)$ & $SO(4, 2)$ extremely useful

- 2 physical modes in

$$\mathcal{H}_{\text{phys}} = \{\mathcal{D}^2 \mathcal{A} = 0, \mathcal{A} \text{ gauge fixed}\} / \{\text{pure gauge}\}$$

$$\mathcal{A}_\mu^{(2+)}[\phi^{(s)}] = \{\mathbf{x}_\mu, \phi^{(s)}\}_+$$

$$\mathcal{A}_\mu^{(2-)}[\phi^{(s)}] = \{\mathbf{x}_\mu, \phi^{(s)}\}_-$$

pure gauge mode

$$\mathcal{A}_\mu^{(g)}[\phi^{(s)}] = \{t_\mu, \phi^{(s)}\}$$

time-like mode $\mathcal{A}_\mu^{(4)}[\phi^{(s)}] = x_\mu \phi^{(s)}$ (not in $\mathcal{H}_{\text{phys}}$)

conjecture: **no ghosts** (cf. YM !)

- all physical on-shell modes found,
2 physical propagating modes for each spin $s \leq n$

further aspects of $\mathcal{M}^{3,1}$:

- symmetry $SO(3, 1)$ only space-like
⇒ "Lorentz-violating" structures: $x_\mu \{t^\mu, \cdot\} \sim \frac{\partial}{\partial \tau}$... time-like VF
(cosmic background!)
- same speed (of light) for all physical modes

vielbein, metric & dynamical geometry

consider scalar field $\phi = \phi(X)$ (e.g. transversal fluctuation)

kinetic term

$$-Tr[Y^a, \phi][Y_a, \phi] \sim \int e^a \phi e_a \phi = \int \gamma^{\mu\nu} \partial_\mu \phi \partial_\nu \phi$$

vielbein

$$\begin{aligned} e^a &:= \{Y^a, \cdot\} = e^{a\mu} \partial_\mu \\ e^{a\mu} &= \theta^{a\mu} \end{aligned}$$

metric

$$\gamma^{\mu\nu} = \eta_{\alpha\beta} e^{\alpha\mu} e^{\beta\nu}$$

dynamical frame bundle, metric

perturbed vielbein: $Y^a = \bar{Y}^a + \mathcal{A}^a$

$$e^a := \{\bar{y}^a, \cdot\} = e^{a\mu}[\mathcal{A}] \partial_\mu \quad \dots \text{vielbein}$$

$$\delta_{\mathcal{A}} \gamma^{ab} =: \{\bar{y}^c, x^a\} \{\mathcal{A}_c, x^b\} + (a \leftrightarrow b)$$

linearize & average over fiber $\rightarrow h^{ab} = [\delta_{\mathcal{A}} \gamma^{ab}]_0$

coupling to matter:

$$S[\text{matter}] \sim \int_{\mathcal{M}} d^4x h^{ab} T_{ab}$$

on H^4 : quadratic action:

$$S_{YM} \propto \int h_{ab} h_{ab}$$

$h_{ab} \sim T_{ab}$, doesn't propagate on H^4 in classical model

HS, arXiv:1606.00769, M. Sperling, HS arXiv:1707.00885

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towards higher-spin gravity on $\mathcal{M}^{3,1}$

perturbed metric: $\delta_{\mathcal{A}} \gamma^{\mu\nu} \propto \{\mathcal{A}^\mu, x^\nu\} + (\mu \leftrightarrow \nu)$

- contains all required dof for gravity !
- **YM action** $\rightarrow 2 \sim$ Ricci-flat grav. waves
+3 propagating modes (cf. massive gravity)
- linearized **Einstein-Hilbert action** obtained (gauge invariance),
expected to be induced by quantum effects (cf. **Sakharov**)
(conformal mode non-standard?)
- YM-action replaces $\int d^4x \sqrt{g}$, **stabilizes** $\mathcal{M}^{3,1}$
no cosm. const. problem ?!
- expect
 - \approx linearized GR
 - significant differences at cosmic scales

summary

- **matrix models:**
natural framework for quantum theory of space-time & matter
- **4D covariant quantum spaces** → higher spin theories
- \exists nice cosmological FRW space-time solutions
reg. BB, finite density of microstates
- Yang-Mills structure → **emergent gravity** rather than GR
- good UV behavior (SUSY), well suited for quantization

... intriguing, needs more work

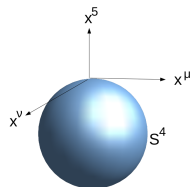
local description: pick north pole $p \in S^4$

→ tangential & radial generators

$$X^a = \begin{pmatrix} X^\mu \\ X^5 \end{pmatrix}, \quad \mu = 1, \dots, 4 \dots \text{tangential coords at } p$$

separate $SO(5)$ into $SO(4)$ & translations

$$\mathcal{M}^{ab} = \begin{pmatrix} \mathcal{M}^{\mu\nu} & \mathcal{P}^\mu \\ -\mathcal{P}^\mu & 0 \end{pmatrix}$$



Poisson algebra $\{P_\mu, X^\nu\} \approx \delta_\mu^\nu$ locally

local form of spin 2 harmonics:

$$\phi^{(2)} = h_{\mu\nu}(x) P^\mu P^\nu + \omega_{\mu;\alpha\beta}(x) P^\mu M^{\alpha\beta} + \Omega_{\alpha\beta;\mu\nu}(x) M^{\alpha\beta} M^{\mu\nu}$$

recall $End(\mathcal{H}) = \oplus \mathcal{C}^s$, $\mathcal{C}^s \cong$ rank s tensor fields $\phi_{a_1 \dots a_s}(x) \cong (n, 2s)$
unique irrep $(n, 2s) \Rightarrow$ **constraints!**

$$\omega_{\mu;\alpha\beta} \propto \partial_\alpha h_{\mu\beta} - \partial_\beta h_{\mu\alpha}$$

$$\Omega_{\alpha\beta;\mu\nu} \propto \mathcal{R}_{\alpha\beta\mu\nu}[h]$$

... linearized spin connection and curvature **determined by** $h_{\mu\nu}$

local spn 1 gauge trafos:

$$\Lambda = \Lambda_{ab}^{(1)}(x)\theta^{ab} = v_\mu(x)P^\mu + \omega_{\mu\nu}(x)\mathcal{M}^{\mu\nu}.$$

$\omega_{\mu\nu}$... field strength of vector field v_μ

$$\delta_\Lambda X^\mu = \{\Lambda, X^\mu\} = - \underbrace{v^\mu}_{\text{diffeo}} + \left(\underbrace{\partial_\nu v_\rho}_{\text{pure gauge graviton}} P^\rho + \partial_\nu \omega_{\rho\sigma} \mathcal{M}^{\rho\sigma} \right) \theta^{\mu\nu} \in \mathcal{C}^0 + \mathcal{C}^2$$